

# CONTRAVARIANT = COVARIANT FOR CARTESIAN VECTORS

43.4

$$x_i' = a_{ij} x_j \Rightarrow a_{ij} = \frac{\partial x_i'}{\partial x_j} \quad \text{old notation} \quad (1)$$

$$x_i' = a_j^i x^j \Rightarrow a_j^i = \frac{\partial x_i'}{\partial x^j} \quad \text{new notation} \quad (2)$$

Invert (2):  $a_i^m x_i' = a_i^m \underbrace{a_j^i}_{\delta_j^m} x^j = x^m \Rightarrow x^m = a_i^m x_i'$

$$\Rightarrow a_i^m = \frac{\partial x^m}{\partial x_i'} \quad (3)$$

In (3) replace  $m$  by  $j$   $\Rightarrow$   $a_i^j = \frac{\partial x^j}{\partial x_i'}$  (4)

Compare this to (2):  $a_j^i = \frac{\partial x_i'}{\partial x^j} \xrightarrow{i \leftrightarrow j} a_i^j = \frac{\partial x^j}{\partial x_i'}$  (5)

Finally, comparing (4) & (5) we see that

$$\frac{\partial x^j}{\partial x_i'} = \frac{\partial x^j}{\partial x_i'} \quad (6)$$

From Eq. (6) we see that it makes no difference whether the primed coordinates are in the numerator or denominator, which is to say that CONTRAVARIANT = COVARIANT

# MANIPULATING TENSORS

45

a) Contraction ("generalized trace")

$$T'^{\mu\rho} = T'^{\mu\rho\nu} \quad (\text{sum over repeated index } \nu) \quad (1)$$

→ Show that this is a 2ND rank tensor

$$T'^{\mu\nu\rho}_{\lambda}(x') = \frac{\partial x'^{\mu}}{\partial x^{\alpha}} \frac{\partial x'^{\rho}}{\partial x^{\beta}} \frac{\partial x'^{\nu}}{\partial x^{\gamma}} \frac{\partial x^{\epsilon}}{\partial x'^{\lambda}} T^{\alpha\beta\gamma}_{\epsilon}(x) \quad (2)$$

Set  $\lambda = \nu$  and sum:

$$T'^{\mu\rho\nu}_{\nu}(x') = \frac{\partial x'^{\mu}}{\partial x^{\alpha}} \frac{\partial x'^{\rho}}{\partial x^{\beta}} \frac{\partial x'^{\nu}}{\partial x^{\gamma}} \frac{\partial x^{\epsilon}}{\partial x'^{\nu}} T^{\alpha\beta\gamma}_{\epsilon}(x) \quad (3)$$

$$\frac{\partial x^{\epsilon}}{\partial x'^{\nu}} \equiv \delta^{\epsilon}_{\gamma} \quad (\text{key step}) \quad (4)$$

Hence:  $T'^{\mu\rho\nu}_{\nu}(x') = \frac{\partial x'^{\mu}}{\partial x^{\alpha}} \frac{\partial x'^{\rho}}{\partial x^{\beta}} T^{\alpha\beta\gamma}_{\gamma}$  (5)

This shows that  $T'^{\mu\rho\nu}_{\nu}(x')$  transforms as a 2ND rank Contravariant tensor.

## b) Raising & Lowering Indices

(45a)

• Define  $S_{\nu\lambda\sigma}^{\mu\rho} \equiv g_{\nu\lambda} T_{\sigma}^{\mu\rho} =$  5TH RANK MIXED TENSOR

• Then set  $\mu = \nu$  & sum  $\Rightarrow S_{\mu\lambda\sigma}^{\mu\rho} = \boxed{S_{\lambda\sigma}^{\rho} = g_{\mu\lambda} T_{\sigma}^{\mu\rho}}$

Hence effect of  $g_{\mu\lambda}$  is to lower an index of the original tensor  $T_{\sigma}^{\mu\rho}$ . This is important when keeping track of indices (an example will follow later - Generalized Laplacian)

• Indices can also be raised using the inverse tensor to  $g_{\mu\nu} \equiv g^{\lambda\mu}$ :

$$g^{\lambda\mu}(x) g_{\mu\nu}(x) = \delta_{\nu}^{\lambda} = \begin{cases} 1 & \mu = \nu \\ 0 & \mu \neq \nu \end{cases}$$

# TENSOR DENSITIES

48

● We have previously defined  $g(x) = \det g_{\mu\nu}(x)$  (1)

Since  $g_{\mu\nu}(x)$  is a 2ND rank covariant tensor it transforms as

$$g'_{\mu\nu}(x') = \frac{\partial x^\rho}{\partial x'^\mu} \frac{\partial x^\sigma}{\partial x'^\nu} g_{\rho\sigma}(x) \quad (2)$$

We can regard  $\partial x^\rho / \partial x'^\mu \equiv f_\mu^\rho$  as the matrix relating  $x'$  and  $x$

in which case Eq. (2) can be written in the form

$$g'_{\mu\nu} = f_\mu^\rho g_{\rho\sigma} f_\nu^\sigma \quad (3)$$

Taking the determinant of both sides of (3) we get

$$\underbrace{\det g'_{\mu\nu}}_{g'} = \det (f_\mu^\rho g_{\rho\sigma} f_\nu^\sigma) = (\det f_\mu^\rho) (\det g_{\rho\sigma}) (\det f_\nu^\sigma) \quad (4)$$

$$\text{where we have used } \det(ABC) = (\det A)(\det B)(\det C) \quad (5)$$

Hence from (4):

$$g'(x') = (\det f)^2 g(x) \equiv \left| \frac{\partial x}{\partial x'} \right|^2 g(x) \quad (6)$$

→ JACOBIAN OF TRANSFORMATION FROM  $x \rightarrow x'$

Since  $g(x)$  is a scalar [it has no indices] the "expected" transformation law would be

$$g'(x') = g(x) \quad (7)$$

Without the additional factor  $(\det f)^2 = \left| \frac{\partial x}{\partial x'} \right|^2$ .

The presence of the additional factor  $|\frac{\partial x}{\partial x'}|^2$  makes  $g(x)$  not just a scalar but a Scalar density.

More generally a quantity which transforms as a tensor except for additional factors of  $|\frac{\partial x'}{\partial x}|^w$  is a tensor density of weight  $w$ .

$$\text{Note that: } \left| \frac{\partial x}{\partial x'} \right| = \left| \frac{\partial x'}{\partial x} \right|^{-1} \quad (8)$$

Combining (8) & (6) we have

$$g'(x') = \left| \frac{\partial x'}{\partial x} \right|^{-2} g(x) \quad (9)$$

Hence  $g(x)$  is a scalar density of weight  $w = -2$ .

We can show that any tensor of weight  $w$  can be expressed as a product of an ordinary tensor multiplied by a factor  $g^{-w/2}$ .

To see this let  $T_{\nu}^{\mu}(x)$  be a tensor density of weight  $w$  so that

$$T'_{\nu}{}^{\mu}(x') = \left| \frac{\partial x'}{\partial x} \right|^w \frac{\partial x'^{\mu}}{\partial x^{\rho}} \frac{\partial x^{\sigma}}{\partial x'^{\nu}} T_{\sigma}^{\rho}(x) \quad (10)$$

$$\therefore [g'^{w/2}] T'_{\nu}{}^{\mu}(x') = \underbrace{\left[ \left| \frac{\partial x'}{\partial x} \right|^{-2} g(x) \right]^{w/2}}_{g(x)^{w/2}} \left| \frac{\partial x'}{\partial x} \right|^w \frac{\partial x'^{\mu}}{\partial x^{\rho}} \frac{\partial x^{\sigma}}{\partial x'^{\nu}} T_{\sigma}^{\rho}(x) \quad (11)$$

$$\therefore \left\{ g'^{w/2} T'_{\nu}{}^{\mu}(x') \right\} = \frac{\partial x'^{\mu}}{\partial x^{\rho}} \frac{\partial x^{\sigma}}{\partial x'^{\nu}} \left\{ g^{w/2}(x) T_{\sigma}^{\rho}(x) \right\} \Rightarrow \quad (12)$$

$$g^{w/2}(x) T_{\sigma}^{\rho}(x) \sim T_{\sigma}^{\rho}(x) \rightarrow \underline{\text{an ordinary tensor}}$$

## ANOTHER SPECIAL TENSOR

50

Levi-Civita tensor  $\epsilon^{\mu\nu\lambda\kappa}$

$$\epsilon^{\mu\nu\lambda\kappa} = \begin{cases} +1 & \text{even permutation of } 1234 \text{ (} x^4_{=ict}) \\ -1 & \text{odd " " " "} \\ 0 & \text{2 indices equal} \end{cases}$$

This symbol plays an important role in formulating Maxwell's equations in 4-dimensional notation.

It allows  $\vec{E}$  &  $\vec{B}$  to be expressed in terms of a single tensor  $F_{\mu\nu}$ .

Notation:

$$F_{\mu\nu} = F_{\nu\mu} \quad (\text{Symmetric tensor})$$
$$F_{\mu\nu} = -F_{\nu\mu} \quad (\text{antisymmetric tensor})$$

Let  $A_{\mu\nu}$  be any tensor. Then

$$A_{\mu\nu} = \underbrace{\frac{1}{2} (A_{\mu\nu} + A_{\nu\mu})}_{\text{SYMMETRIC}} + \underbrace{\frac{1}{2} (A_{\mu\nu} - A_{\nu\mu})}_{\text{ANTISYMMETRIC}}$$

# Transformation of the Levi-Civita Symbol $\epsilon^{M\nu\lambda k}$ :

50/51

Recall:  $\epsilon^{1234} = +1$  = even permutations of 1234

$\epsilon^{2134} = -1$  = even permutations of 2134  
= odd permutations of 1234

$\epsilon^{M\nu\lambda k} = 0$  when 2 indices are equal

To show that  $\epsilon^{M\nu\lambda k}$  is a tensor density consider the expression

$$\sum \rho^{\sigma\eta\xi} \equiv \frac{\partial x'^{\rho}}{\partial x^{\mu}} \frac{\partial x'^{\sigma}}{\partial x^{\nu}} \frac{\partial x'^{\eta}}{\partial x^{\lambda}} \frac{\partial x'^{\xi}}{\partial x^k} \epsilon^{M\nu\lambda k} \quad (1)$$

We can see that  $\sum \rho^{\sigma\eta\xi}$  must be proportional to  $\epsilon^{\rho\sigma\eta\xi}$  by noting that (for example)  $\sum \rho^{\rho\rho\xi} = 0$ . This follows immediately by noting from (1) that setting  $\rho = \sigma$  on the r.h.s. of (1) has the effect of making the coefficient of  $\epsilon^{M\nu\lambda k}$  SYMMETRIC in  $\mu \leftrightarrow \nu$ , whereas  $\epsilon^{M\nu\lambda k}$  itself is ANTISYMMETRIC, and hence their product vanishes. We can thus write

$$\sum \rho^{\sigma\eta\xi} = C \cdot \epsilon^{\rho\sigma\eta\xi} \quad (2)$$

$\uparrow$   
constant

To evaluate C write:

$$\sum 1^{234} \equiv C \epsilon^{1234} = \frac{\partial x'^1}{\partial x^1} \frac{\partial x'^2}{\partial x^2} \frac{\partial x'^3}{\partial x^3} \frac{\partial x'^4}{\partial x^4} \epsilon^{1234} \quad (3)$$

$\swarrow +1$

$$+ \frac{\partial x'^2}{\partial x^2} \frac{\partial x'^1}{\partial x^1} \frac{\partial x'^3}{\partial x^3} \frac{\partial x'^4}{\partial x^4} \epsilon^{2134} + \dots$$

$\swarrow -1$   
all remaining permutations of 1234

It can be seen that the expression on the r.h.s. of (3) is just the determinant of the transformation from  $x \rightarrow x'$ !

$$\text{r.h.s.} = \begin{vmatrix} \frac{\partial x'^1}{\partial x^1} & \frac{\partial x'^2}{\partial x^1} & \frac{\partial x'^3}{\partial x^1} & \frac{\partial x'^4}{\partial x^1} \\ \frac{\partial x'^1}{\partial x^2} & \frac{\partial x'^2}{\partial x^2} & \frac{\partial x'^3}{\partial x^2} & \frac{\partial x'^4}{\partial x^2} \\ \dots & \dots & \frac{\partial x'^3}{\partial x^3} & \frac{\partial x'^4}{\partial x^3} \\ \dots & \dots & \dots & \frac{\partial x'^4}{\partial x^4} \end{vmatrix} \quad (4)$$

From (3) & (4) we thus see that

$$\underline{X}^{1234} = c \underbrace{\epsilon'^{1234}}_{+1} = \left| \frac{\partial x'}{\partial x} \right| \cdot 1 \Rightarrow c = \left| \frac{\partial x'}{\partial x} \right| \quad (5)$$

Combining the previous results we find

$$\underline{X}^{\rho\sigma\eta\xi} = \left| \frac{\partial x'}{\partial x} \right| \epsilon'^{\rho\sigma\eta\xi} = \frac{\partial x'^\rho}{\partial x^\mu} \frac{\partial x'^\sigma}{\partial x^\nu} \frac{\partial x'^\eta}{\partial x^\alpha} \frac{\partial x'^\xi}{\partial x^\kappa} \epsilon^{\mu\nu\alpha\kappa} \quad (6)$$

Hence

$$\epsilon^{\rho\sigma\eta\xi} = \left| \frac{\partial x'}{\partial x} \right|^{-1} \frac{\partial x'^\rho}{\partial x^\mu} \frac{\partial x'^\sigma}{\partial x^\nu} \frac{\partial x'^\eta}{\partial x^\alpha} \frac{\partial x'^\xi}{\partial x^\kappa} \epsilon^{\mu\nu\alpha\kappa} \quad (7)$$

This establishes that  $\underline{\epsilon}^{\mu\nu\alpha\kappa}$  is a 4th rank tensor density of weight  $W = -1$



# THE AFFINE CONNECTION [See WEINBERG "Gravitation & Cosmology"]

53  
53.1  
61

The affine connection  $\Gamma_{\mu\nu}^{\lambda}(x)$  is NOT a tensor. In fact its importance derives precisely from the fact that this is so.

The affine connection is part of the formalism for COVARIANT

## DIFFERENTIATION:

$$\underbrace{\frac{\partial}{\partial x^{\lambda}} V^{\mu}(x)}_{\text{not a tensor}} \longrightarrow \underbrace{\frac{\partial}{\partial x^{\lambda}} V^{\mu}(x)}_{\text{not a tensor}} + \underbrace{\Gamma_{\lambda\nu}^{\mu} V^{\nu}(x)}_{\text{not a tensor}} = \underbrace{D_{\lambda} V^{\mu}(x)}_{\text{a tensor}} \quad (1)$$

We show that the "bad" parts of  $\frac{\partial}{\partial x^{\lambda}} V^{\mu}$  which make it not a tensor, exactly cancel against the "bad" parts of  $\Gamma_{\lambda\nu}^{\mu} V^{\nu}$  so that their sum is indeed a tensor.

## MOTIVATING THE AFFINE CONNECTION:

From elementary physics we know that (Newton 2<sup>nd</sup> Law)

$$\begin{array}{l} \text{no forces} \Rightarrow \underbrace{\frac{d^2 \vec{x}}{dt^2}}_{\text{3-dim}} \equiv 0 \longrightarrow \underbrace{\frac{d^2 \xi^{\alpha}}{d\tau^2}}_{\text{3-dim}} \quad d\tau \leftrightarrow dt \quad (2) \\ \downarrow \\ \text{"inertial"} \\ \text{Coordinate system} \end{array}$$

$$\hookrightarrow c^2 d\tau^2 = c^2 dt^2 - (d\vec{x})^2 = -\eta_{\alpha\beta} d\xi^{\alpha} d\xi^{\beta}$$

QUESTION: What does "Newton II" look like in a non-inertial coordinate system?

ANSWER: New "forces" arise (centrifugal, Coriolis ...)

The affine connection describes these new "forces"

Define

$$\xi^\alpha = \xi^\alpha(x^\mu) \quad \text{or} \quad x^\mu = x^\mu(\xi^\alpha) \quad (3)$$

$\uparrow$  inertial       $\uparrow$  non-inertial

$$(2) \Rightarrow \frac{d}{d\tau} \left( \frac{\partial \xi^\alpha}{\partial t} \right) = 0 = \frac{d}{d\tau} \left( \frac{\partial \xi^\alpha}{\partial x^\mu} \cdot \frac{dx^\mu}{d\tau} \right) \quad (4)$$

introduces new coord system

$$= \left( \frac{d}{d\tau} \frac{\partial \xi^\alpha}{\partial x^\mu} \right) \left( \frac{dx^\mu}{d\tau} \right) + \frac{\partial \xi^\alpha}{\partial x^\mu} \frac{d^2 x^\mu}{d\tau^2} \rightarrow \text{What we want: acceleration in new coord system} \quad (5)$$

$$\frac{\partial}{\partial x^\mu} \frac{d \xi^\alpha}{d\tau} = \frac{\partial}{\partial x^\mu} \left( \frac{\partial \xi^\alpha}{\partial x^\nu} \frac{dx^\nu}{d\tau} \right) = \frac{\partial^2 \xi^\alpha}{\partial x^\mu \partial x^\nu} \frac{dx^\nu}{d\tau} + \frac{\partial \xi^\alpha}{\partial x^\nu} \frac{\partial}{\partial x^\mu} \left( \frac{dx^\nu}{d\tau} \right) \quad (6)$$

$$\frac{d}{d\tau} \left( \frac{\partial x^\nu}{\partial x^\mu} \right) = 0$$

$\underbrace{\hspace{2cm}}_{\delta^\nu_\mu}$

Altogether: (4)-(6)  $\Rightarrow$

$$0 = \frac{\partial^2 \xi^\alpha}{\partial x^\mu \partial x^\nu} \left( \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right) + \frac{\partial \xi^\alpha}{\partial x^\mu} \frac{d^2 x^\mu}{d\tau^2} \rightarrow \text{What we want} \quad (7)$$

Multiply each term in this equation by  $\partial x^\lambda / \partial \xi^\alpha$  using

$$\frac{\partial \xi^\alpha}{\partial x^\mu} \frac{\partial x^\lambda}{\partial \xi^\alpha} = \frac{\partial x^\lambda}{\partial x^\mu} = \delta^\lambda_\mu \Rightarrow \quad (8)$$

$$0 = \frac{\partial^2 \xi^\alpha}{\partial x^\mu \partial x^\nu} \frac{\partial x^\lambda}{\partial \xi^\alpha} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + \delta^\lambda_\mu \frac{d^2 x^\mu}{d\tau^2}$$

$$0 = \frac{d^2 x^\lambda}{d\tau^2} + \left( \frac{\partial x^\lambda}{\partial \xi^\alpha} \frac{\partial^2 \xi^\alpha}{\partial x^\mu \partial x^\nu} \right) \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau}$$

$\hookrightarrow \Gamma^\lambda_{\mu\nu} = \text{AFFINE CONNECTION}$

(9)

Eq. (9) can be written in the form:

$$\boxed{\frac{d^2 x^\lambda}{d\tau^2} = - \Gamma_{\mu\nu}^\lambda \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau}} \quad (10)$$

Compare this to the usual formulation of Newton II

$$\frac{d^2 \vec{x}}{dt^2} = \frac{\vec{F}}{m} \quad \Leftrightarrow \quad \boxed{\frac{d^2 x^i}{dt^2} = \frac{F^i}{m}} \quad (11)$$

We see that  $\Gamma_{\mu\nu}^\lambda$  plays the role of the external forces acting on a system: Later we will see that since  $\Gamma_{\mu\nu}^\lambda$  enters into the formula for the covariant derivative,

We can interpret covariant differentiation as introducing forces in a natural way into a free particle Hamiltonian or Lagrangian. This is one reason why  $\Gamma_{\mu\nu}^\lambda$  is important.

### Relation Between $\Gamma_{\mu\nu}^\lambda$ and $g_{\mu\nu}$ :

Start with an earlier equation:

$$g_{\mu\nu}(x) = \frac{\partial \xi^\alpha}{\partial x^\mu} \frac{\partial \xi^\beta}{\partial x^\nu} \eta_{\alpha\beta}(\xi) \quad \left( \begin{matrix} 1 & 0 \\ 0 & 1 \end{matrix} \right) \quad x_4^4 = ict \quad (12)$$

$$\frac{\partial g_{\mu\nu}}{\partial x^\lambda} = \frac{\partial^2 \xi^\alpha}{\partial x^\lambda \partial x^\mu} \frac{\partial \xi^\beta}{\partial x^\nu} \eta_{\alpha\beta} + \frac{\partial \xi^\alpha}{\partial x^\mu} \frac{\partial^2 \xi^\beta}{\partial x^\lambda \partial x^\nu} \eta_{\alpha\beta} \quad (13)$$

$\swarrow \quad \searrow$ 
 $\Gamma_{\lambda\mu}^\rho \frac{\partial \xi^\alpha}{\partial x^\rho}$ 
 $\swarrow \quad \searrow$ 
 $\Gamma_{\lambda\nu}^\rho \frac{\partial \xi^\beta}{\partial x^\rho}$

Hence  $\frac{\partial g_{\mu\alpha}}{\partial x^\lambda} = \Gamma_{\lambda\mu}^\rho \left( \frac{\partial \xi^\alpha}{\partial x^\rho} \frac{\partial \xi^\beta}{\partial x^\nu} \eta_{\alpha\beta} \right) \rightarrow g_{\rho\nu}$  (14) 55

$$+ \Gamma_{\lambda\nu}^\rho \left( \frac{\partial \xi^\alpha}{\partial x^\mu} \frac{\partial \xi^\beta}{\partial x^\rho} \eta_{\alpha\beta} \right) \rightarrow g_{\mu\rho}$$
 (15)

$$\therefore \frac{\partial g_{\mu\nu}}{\partial x^\lambda} = \Gamma_{\lambda\mu}^\rho g_{\rho\nu} + \Gamma_{\lambda\nu}^\rho g_{\mu\rho}$$
 (16)

For H.W.: You will show that proceeding similarly,

$$\frac{\partial g_{\mu\nu}}{\partial x^\lambda} + \frac{\partial g_{\lambda\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\lambda}}{\partial x^\nu} = 2 g_{\rho\nu} \Gamma_{\lambda\mu}^\rho$$
 (17)

Hence finally, using  $g^{\nu\sigma} g_{\rho\nu} = \delta_\rho^\sigma$  (18)

"METRIC COMPATIBILITY CONDITION"

$$\Gamma_{\lambda\mu}^\sigma = \frac{1}{2} g^{\nu\sigma} \left( \frac{\partial g_{\mu\nu}}{\partial x^\lambda} + \frac{\partial g_{\lambda\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\lambda}}{\partial x^\nu} \right)$$
 (19)

↑ not a tensor    
 ↑ a tensor    
 ↑ not tensors    
 ↑ not tensors

NOTES: (a)  $\Gamma_{\lambda\mu}^\sigma = \Gamma_{\mu\lambda}^\sigma$  [See Eq. (19) above] (10)

(b) In an inertial coord system

$$g_{\mu\nu} \rightarrow \eta_{\mu\nu} = \begin{pmatrix} 1 & & \\ & 1 & \\ & & \dots \end{pmatrix}$$
 (11)

$$\Rightarrow \Gamma_{\lambda\mu}^\sigma \equiv 0$$